

# SSA22 天域の高赤方偏移 4 次元探査

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for Subaru 2005 Intensive Group

初代星・銀河形成研究会  
2008 年 9 月 10 日 甲南大学

## Subaru 2005 Intensive Group ;

山田 亨 (PI), 林野 友紀, (山内 良亮), 中村 有希, 香西 克紀, 森本 奈々,  
中村 江里, 堀江 光典 (東北大), 松田 有一 (国立天文台), 梅村 雅之 (筑波大)

- 2002 ~ 2004 ~ 2005 ~
- SSA22  $z=3.1$  Deep NB Survey for LAE/LAB/LAAs (SCam 7Fovs)
- $V_{com} \approx 200 \times 100 \times 60 \text{Mpc}^3$  (SSA22) ;  $\delta_{200} \approx 1$   
(control field=SXDF 3FoV+SDF)

Huge Over-density Region !

- Standard CDM S.F.(ex. Millennium Simulation) ~  $5\sigma$  event  
==> Cosmology (initial fluctuation?, BAO?, at large/small? scale, CMB)  
==> Very Early Galaxy Formation (GRB/HN/SN at  $z \gg 10$ ?)
- Also, SSA22 Huge ODR ~ beyond 200Mpc scale.

=====> SCALE UP !

- SCam ; by factor(Intensive $\times 2 \sim 3$ )
- HSC ; by order( $> \times 10$ )

# 研究の発端

## Protocluster region at high redshift

Steidel et al. (1998; 2000)

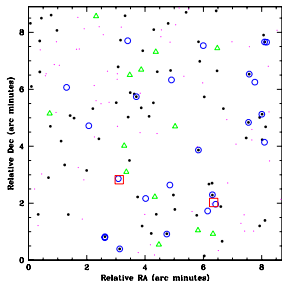
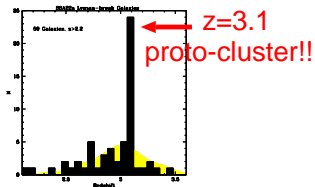
14 天域で  $z=2.7\sim 3.4$  LBG を狙った pencil-beam survey ( $9' \times 9'$  FoV)

- SSA22 天域 (RA= $22^h 17^m$ , Dec= $+00^\circ 15'$ ) の  $z=3.1$  に LBG 高密度領域を発見

⇒ 期待値の約 6 倍の数密度

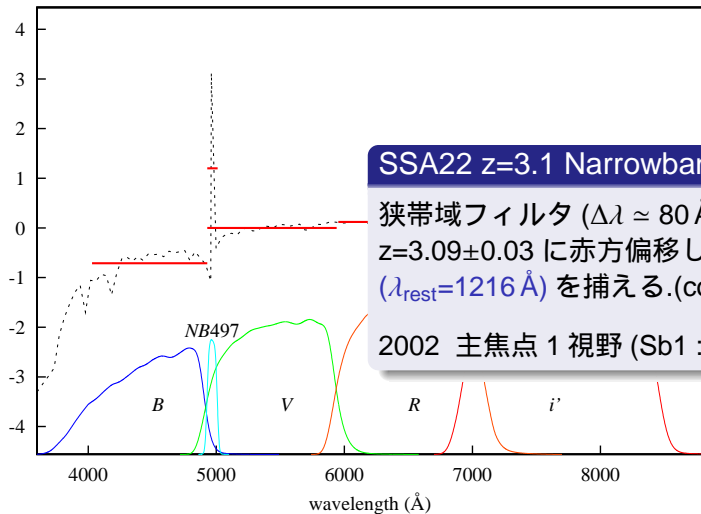
$z=3.09 \pm 0.03$  LAE を狙った Narrowband (NB) survey

- 72 LAEs ( $NB < 25.5\text{Å}$ ,  $EW_0 > 20\text{Å}$ )  
⇒  $9' \times 9'$  FoV 内に一様に分布
- LAE 密度も期待値の約 6 倍



Steidel's survey area ;  $9' \times 9'$  ( 15Mpc $\times$ 15Mpc com. at  $z=3$  )

# SSA22 $z=3.09$ LAE 二次元探査



# SUBARU Extensive Narrowband Deep Imaging Survey

2005 Intensive Program

PI: 山田 亨 (東北大学)

17 nights  $\Rightarrow$  内 10 夜を撮像

## SSA22 $z=3.09$ Extensive Deep Imaging

Sb1+Sb2 ( $\approx 0.43 \text{ deg}^2$ , Hayashino et al. 2004; Yamauchi et al. 2006) のさらなる拡張隣接探査

$\Rightarrow$  SuprimeCam 5 視野分の隣接領域を撮像 (2005/8~10)

各視野同じ深さで観測：

- NB497 (5.5 hours), B (1.0 hours), V (1.0 hours)
- Seeing FWHM  $\approx 1.0 \text{ arcsec}$

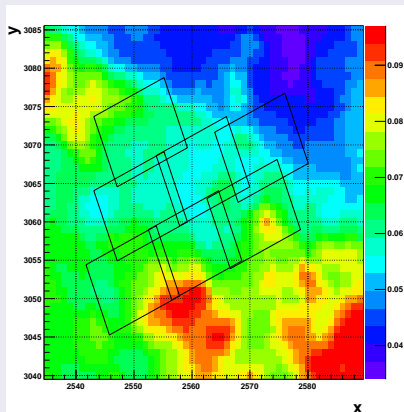
更に一般領域として SXDF(C,N,S), SDF, GOODS-N の 3 天域 (計 5 視野) に対し、同程度の深い NB497 撮像. ( to see SSA22 overdensity  $\delta$  )

# 銀河系吸収補正

## $E(B - V)$ map around SSA22

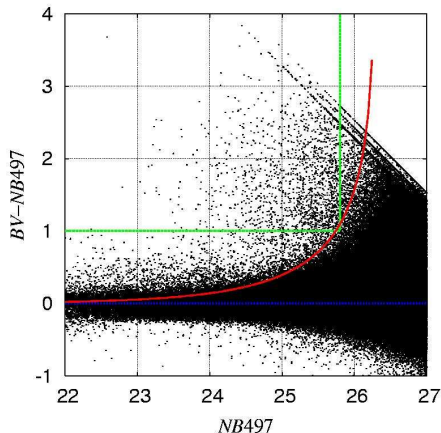
Schlegel et al. (1998)

COBE 100 $\mu$ m-MAP



- SSA22 の銀緯 :  $-40^\circ$  ; 低い
- Schlegel et al. (1998) : 銀河系の有意な吸収量
- Cardelli et al. (1989) による銀河系吸収曲線
- NB497 で 0.2 ~ 0.4 等の吸収量 : ローカルな補正が必要

# Detection criteria for SSA22 $z=3.09\pm 0.03$ LAE



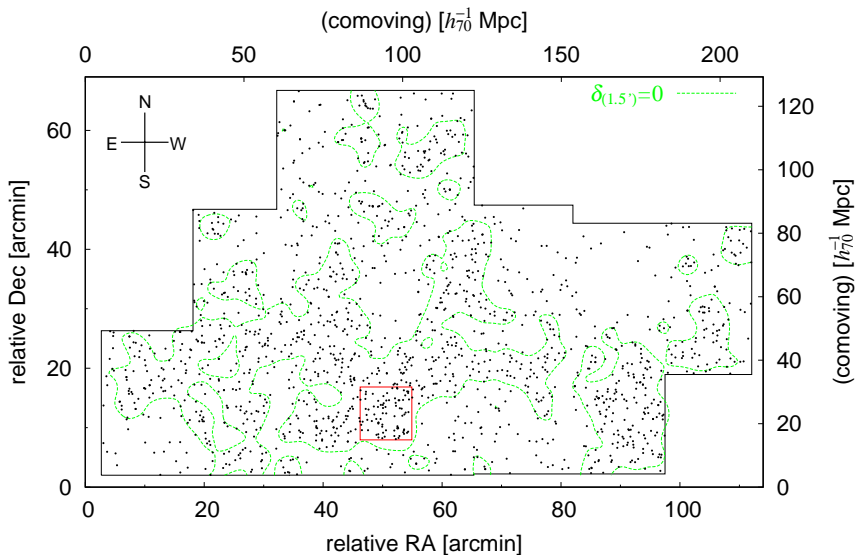
continuum:  $BV = (2B + V)/3$

## LAE 候補の検出

- (1)  $NB497 < 25.8_{AB}$   
( $\Leftrightarrow Ly\alpha$  光度 :  
 $1.3 \times 10^{-17} \text{ erg/s/cm}^2$  以上)
- (2)  $BV - NB497 > 1.0$   
( $\Leftrightarrow EW_0 > 28 \text{ \AA}$ )
- (3)  $\sigma_{(BV-NB)} > 4.0$  (for Sb2 field)
- (4)  $V_c > 4\sigma$ :  $B - V_c > 0.2$
- (5)  $V_c < 4\sigma$ :  $BV - NB497 > 1.3$

$\Rightarrow 1.3 \text{ deg}^2$  ( $\approx 4700$  平方分) の探査領域において、  
1598 個の LAE 候補を検出

# 2D Sky Map of LAEs in SSA22 $z=3.09\pm 0.03$



Total Survey Volume  $\approx 200\times 100\times 60$  Mpc<sup>3</sup>



# Overdensity of LAEs

Control fields: SXDF (C,N,S)+SDF

Overdensity:

SXDF (C,N,S):	$0.179 \text{ } \square'^{-2}$	424 個/2363 $\square'^2$
SDF:	$0.245 \text{ } \square'^{-2}$	154 個/628 $\square'^2$

$$\delta = \frac{\rho - \langle \rho \rangle}{\langle \rho \rangle}$$

$$\Rightarrow \langle \rho \rangle = 0.193 \text{ } \square'^{-2}$$

Region	Area [arcmin <sup>2</sup> ]	$N_{LAE}$	$\delta_{LAE}$	$\sigma_m$	$\sigma_{gal}$
SSA22 (Steidel)	100	100	$4.18 \pm 0.29$	0.19	0.57
SSA22 (sb1)	768	364	$1.46 \pm 0.14$	0.10	0.30
SSA22 (entire)	4726	1598	$0.75 \pm 0.09$	0.055	0.17
Control fields	2991	578	0	0.064	0.19

$\sigma_{gal} = b \times \sigma_m$ , here,  $b=3$ , 60Mpc thick each

SSA22(entire) の出現確率は  $4 \sim 5\sigma (10^{-5 \sim -6})$  : Huge over-density region

## Overdensity $\delta$ of High $EW_0$ LAEs

NB497 < 25.8, (< 25.4)

BV-NB = 1.0~1.7, 1.7~2.2, > 2.2

BV-NB      corrected  $EW_0$

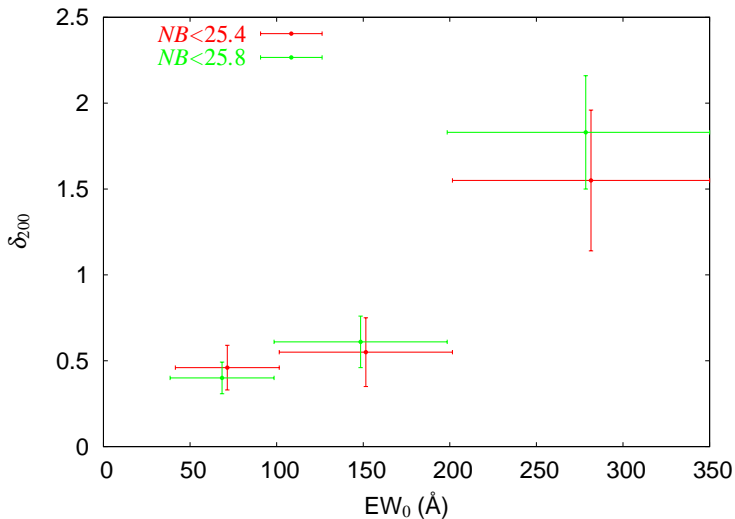
1.0             $\approx 40\text{\AA}$

1.7             $\approx 100\text{\AA}$

2.2             $\approx 200\text{\AA}$

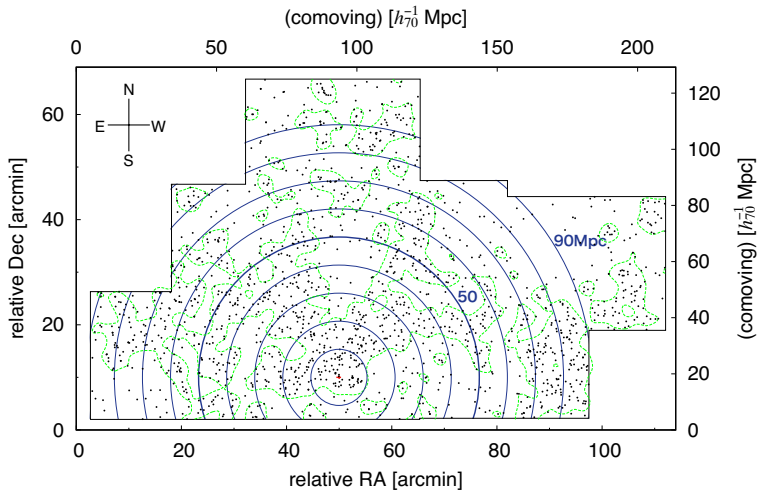
- $\delta_{200}(200 \times 100 \times 60 Mpc^3)$  vs  $EW_0$
- $\delta_{ring}(R)$  ; radial profile of over-densities around Steidel's density peak
- (Anti-) correlation between Very Young(High  $EW_0$ ) LAE & LBG-like LAE(BV-NB=1.0-1.2,  $V < 26$ )
- $Ly\alpha$  Size(NB497 FWHM) as a function of  $EW_0$

# $\delta_{200}$ vs $EW_0$

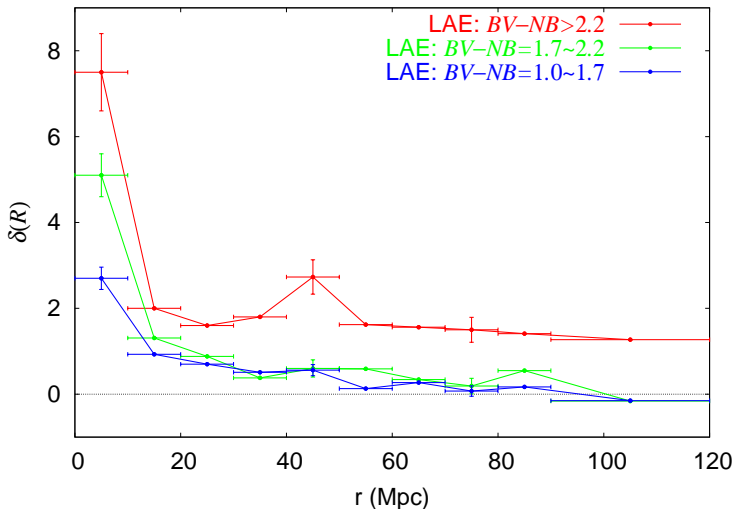


# Radial Profile of Over-density $\delta_{ring}(R)$ around Steidel's density peak

$z=3.1$  LAE( $EW_{0,cor} \geq 40 \text{ \AA}$ ) sky map

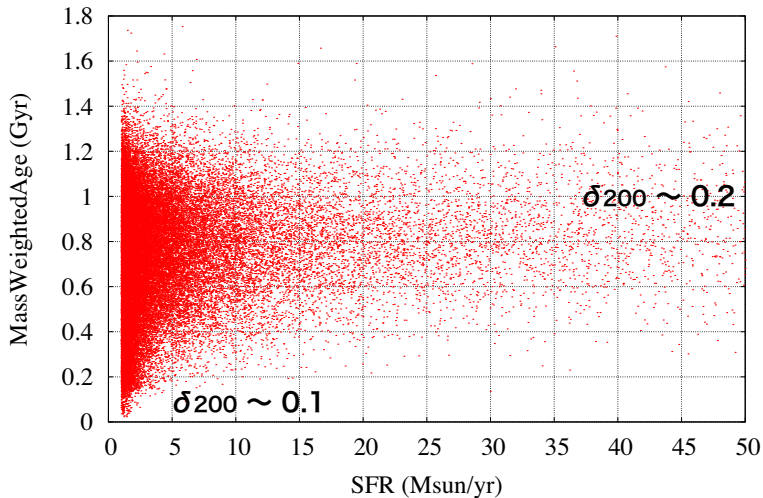


$$\delta_{ring}(R) \quad (1)$$

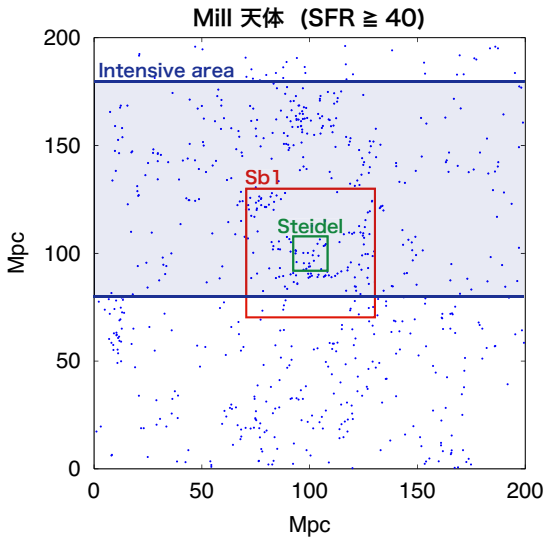


# Millennium Simulation Galaxies ; SFR vs Age

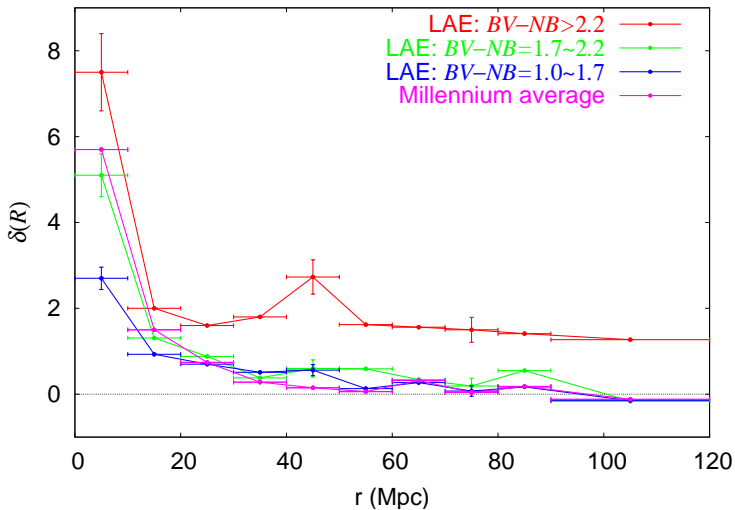
SFR\_MassWeightedAge (Millennium\_sim z=3.1)



# SSA22 like(?) Sky Map of Mill. Galaxies(SFR > 40M<sub>s</sub>/y)



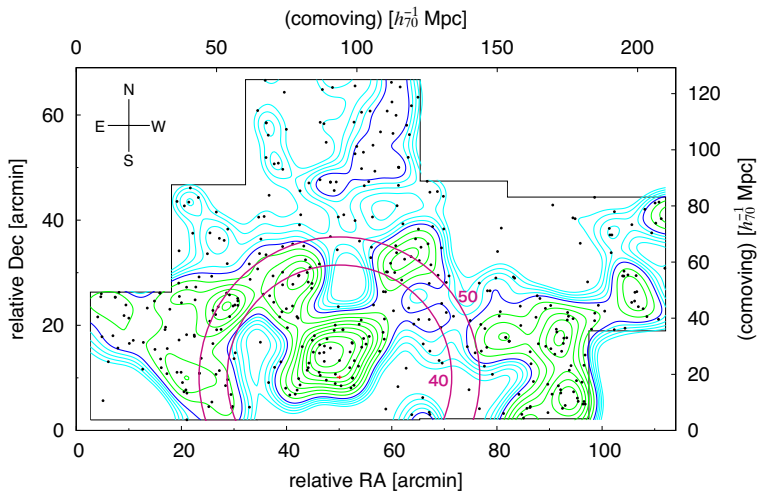
$$\delta_{ring}(R) \quad (2)$$



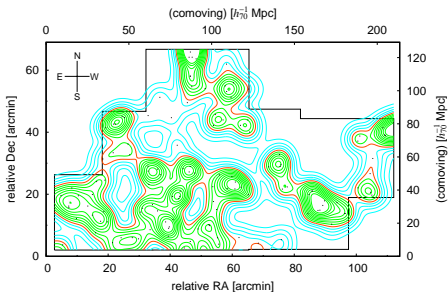
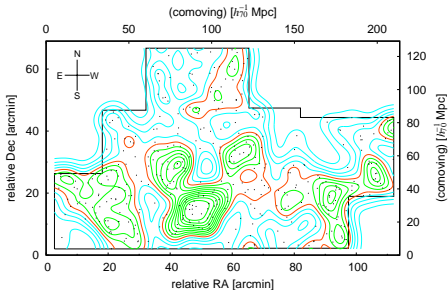


# Ring like LSS? of High $EW_0$ LAEs

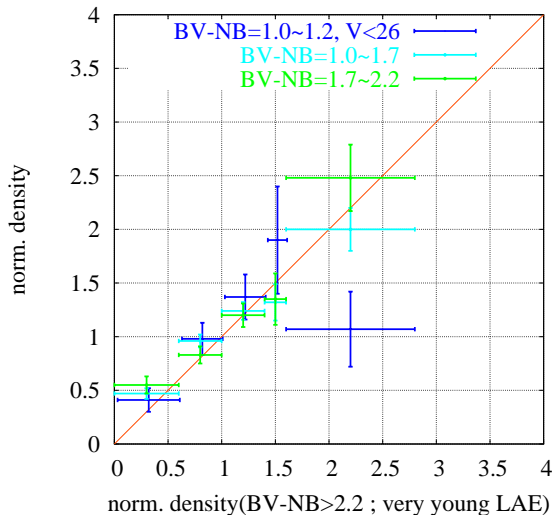
green ; high density region  
sky blue ; low density region



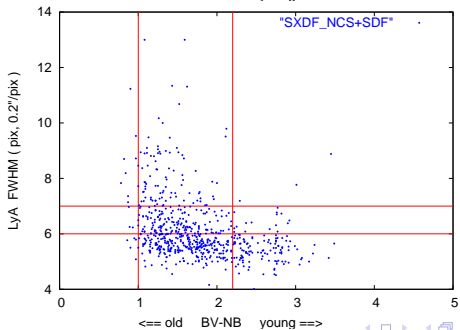
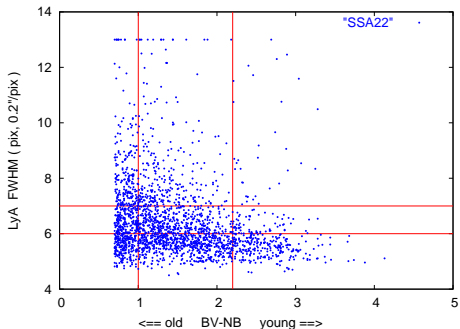
# SKY MAPS for Very Young LAEs and LBG-like LAEs



# Anti-correlation between Very Young LAE & LBG-like LAE (BV-NB=1.0-1.2, V<26)



# Ly $\alpha$ Size(fwhm) vs EW $_0$ (BV-NB)



$$\delta_{200}(EW_{0,cor} > 200\text{\AA} \ \& \ fwhm > 1.2'') \approx 3 !$$

$$\delta_{200}(EW_{0,cor} > 200\text{\AA} \ \& \ fwhm > 1.4'') \approx 6 !!$$

SSA22  $z \sim 3$  Narrowband Imaging Survey

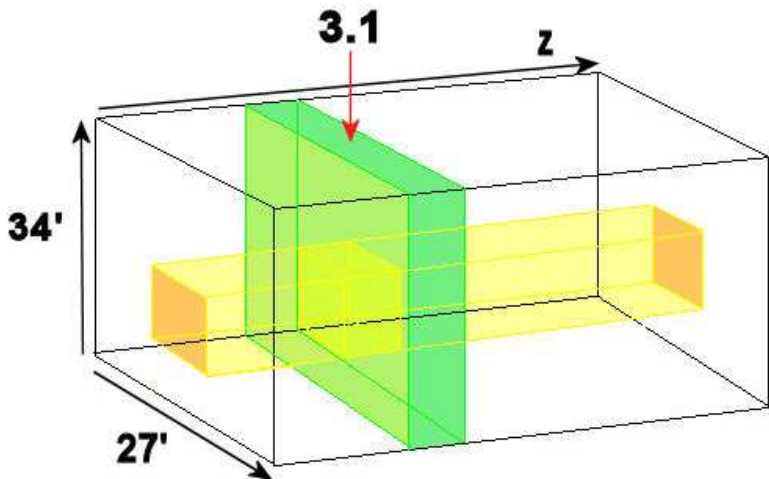
=====> need SCALE UP !

SCam ; by factor(Intensive $\times 2 \sim 3$ ), soon!

====> HSC ; by order(Intensive $\times 10$ ) and  
Survey for Other LSSs/Huge-ODRs

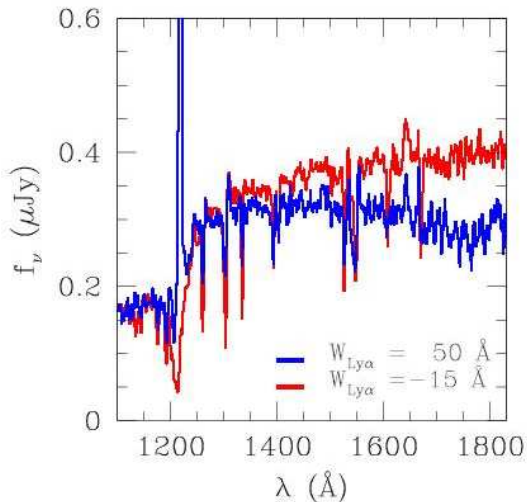
# SSA22-Sb1 $z \sim 3$ LBG 三次元探查

VLT VIMOS obs. 2006 by Yamauchi et al.



# LBG Composite Spectra by Steidel et al.

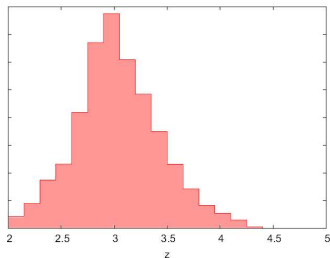
blue ; emLBG  
red ; absLBG





# LBG Window Functions

Ours



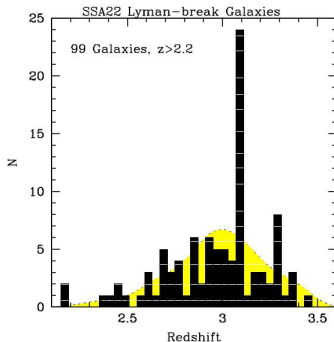
$$\text{WF}(3.3) / \text{WF}(3.1) = 0.65$$

$$V < 25.5$$

$$(U-V) - 1.8(V-R) > 0.85$$

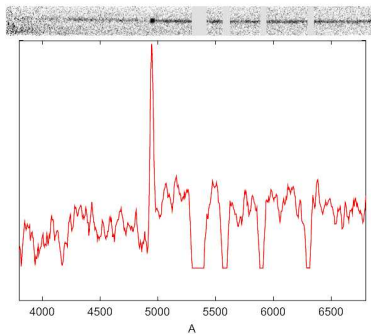
$$R-i < 0.35$$

Steidel's SSA22

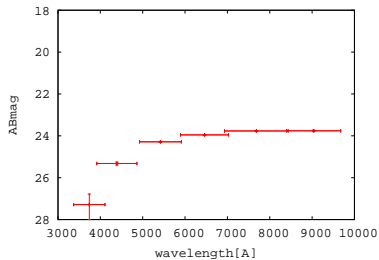
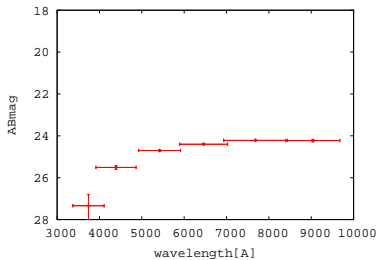
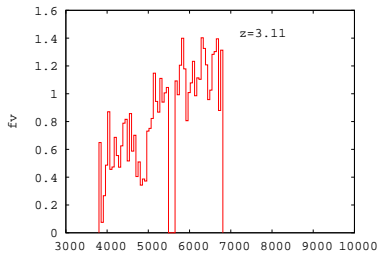
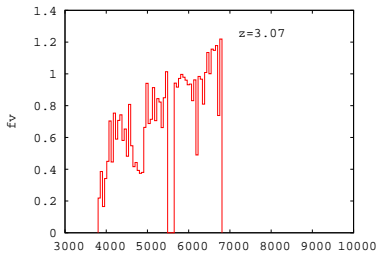


$$\text{WF}(3.3) / \text{WF}(3.1) = 0.4$$

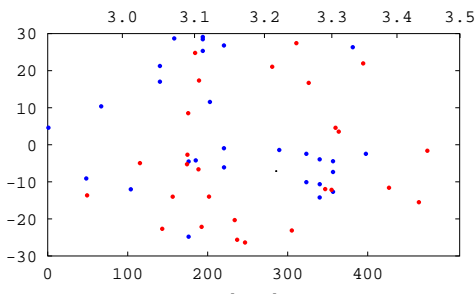
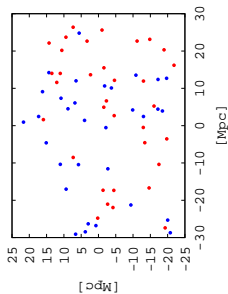
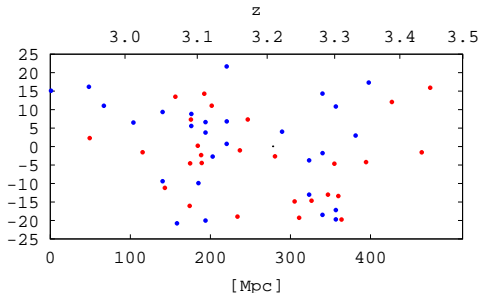
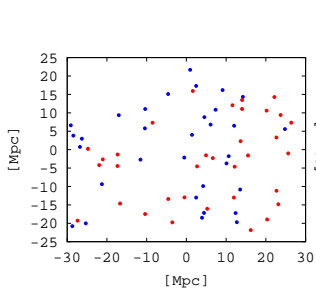
## VIMOS emLBG (example)



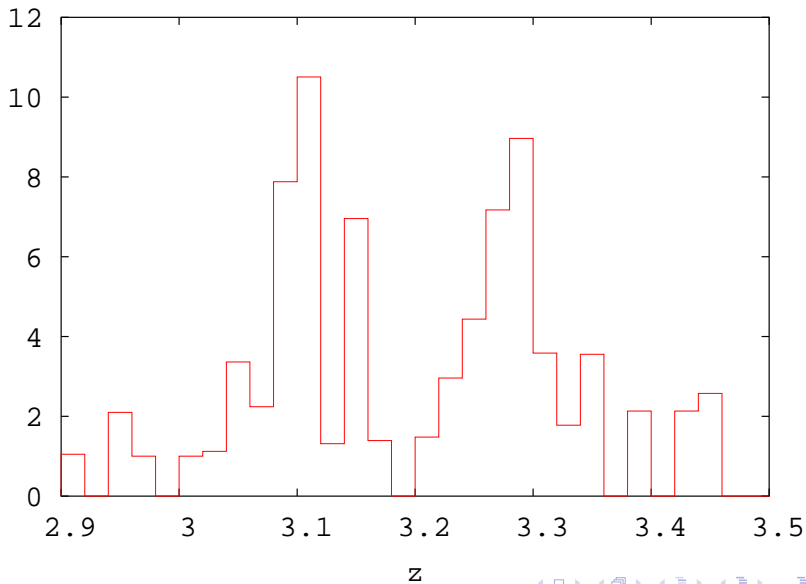
# VIMOS absLBGs (example)



# SSA22-Sb1 LBG 三面図 (red=emLBG, blue=absLBG)



# LBG(emLBG+absLBG) Redshift Distribution



A. Inoue VIMOS 2008 Obs.  
==> More(~3 times) Statistics!!

# SSA22-Sb1 z~3 四次元 (目) 探査 = 変光探査

SSA22 Sb1

2002 B(1.8h), V(1.2h) (ours)

2003 B(0.8h), V(0.8h) (Capak archived)

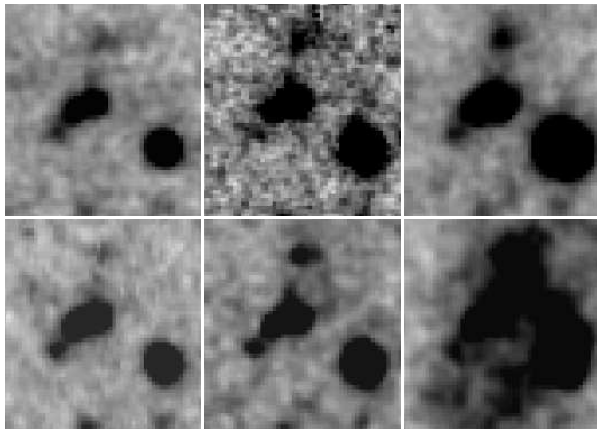
# Candidate Time Variability in LAB2

B02

V02

R02

27.1mag(422adu)



B03

V03

NB497(Ly $\alpha$ )

26.5mag(685adu)

V :  $\Delta\text{flux}(03-02)=263\text{adu}$  ( $1\sigma=120\text{adu}$ )